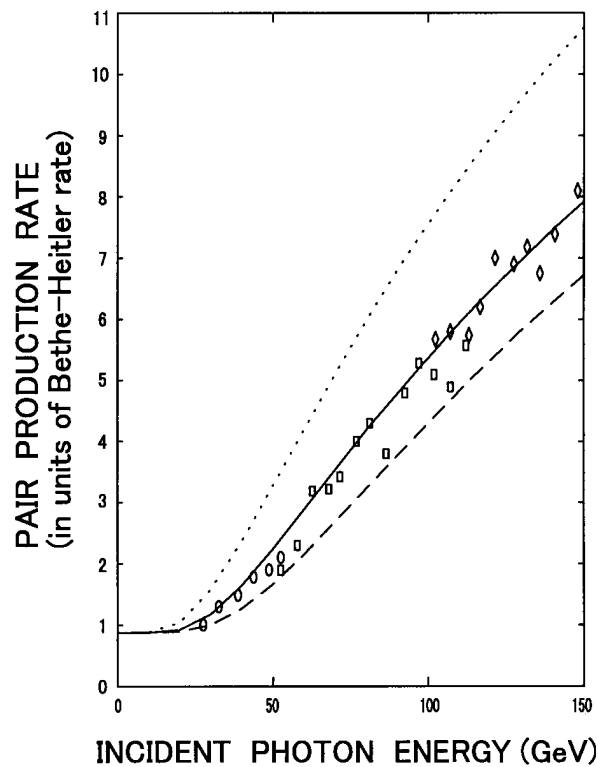


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Semiclassical Theory of Crystal-Assisted Pair Production: Beyond the Uniform Field Approximation

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【Classical Radiation】

Schwinger(Phys.Rev.75 1912(1949))

$$\frac{dN}{d\omega}(t) = \frac{2\alpha}{\pi\gamma^2} J(t, \omega)$$

$$J = \int_0^\infty \gamma^2 (1 - \boldsymbol{\beta}_+ \cdot \boldsymbol{\beta}_-) \sin \Delta_s \frac{d\tau}{\tau} - \frac{\pi}{2} \quad (1)$$

$$\Delta_s = \omega(\tau - |\mathbf{r}_+ - \mathbf{r}_-|/c)$$

where α is the fine structure factor, γ the Lorentz factor, $\mathbf{r}_\pm, \boldsymbol{\beta}_\pm$ the position and velocity at time $t \pm \tau/2$, respectively:

$$\begin{cases} t_+ = t + \tau/2 \\ t_- = t - \tau/2 \end{cases} \quad \begin{cases} \boldsymbol{\beta}_+ = \boldsymbol{\beta}(t_+) \\ \boldsymbol{\beta}_- = \boldsymbol{\beta}(t_-) \end{cases} \quad \begin{cases} \mathbf{r}_+ = \mathbf{r}(t_+) \\ \mathbf{r}_- = \mathbf{r}(t_-) \end{cases}$$

Decomposing the motion of the electron as longitudinal and transverse ones.

$$\begin{cases} \mathbf{r}(t) = \boldsymbol{\rho}(t) + z(t) \\ \boldsymbol{\beta}(t) = \boldsymbol{\beta}_\perp(t) + \beta_z(t) \end{cases}$$

Eq.(1) may be rewritten in the form that includes only the transverse motion:

$$J = \int_0^\infty \left[1 + \frac{\gamma^2}{2} (\delta\beta_\perp)^2 \right] \sin \Delta_s \frac{d\tau}{\tau} - \frac{\pi}{2}$$

$$\Delta_s = \frac{\omega\tau}{2\gamma} - \frac{\omega}{2c^2\tau} (\delta\rho)^2 + \frac{\omega}{2} \int_{t_-}^{t_+} \beta_\perp^2(\tau') d\tau' \quad (2)$$

$$\begin{cases} \delta\beta_\perp = \beta_\perp(t_+) - \beta_\perp(t_-) \\ \delta\rho = \rho(t_+) - \rho(t_-) \end{cases}$$

By expanding factors in terms of τ , we obtain,

$$\Delta_s = \frac{3}{2}\xi(x + x^3/3 + a_5x^5 + a_7x^7 + \dots)$$

$$1 + \frac{\gamma^2}{2}(\delta\beta_\perp)^2 = (1 + 2x^2 + b_4x^4 + b_6x^6 + \dots)$$

where x and ξ are Schwinger's dimensionless time and frequency parameters, respectively:

$$x = \frac{g\gamma\tau}{2}, \quad \xi = \frac{2\omega}{3g\gamma^3}.$$

g the acceleration divided by c . $a_5, a_7, \dots, b_4, b_6, \dots$ may be represented by the time derivatives of g and γ .

【Synchrotron Approximation】

If we approximate the instantaneous trajectory about $t = 0$ by a circular path, then we have $a_5, b_4 \sim 1/\gamma^2$, $a_7, b_6 \sim 1/\gamma^3 \dots$. So, by neglecting the higher order terms under the condition $\gamma \gg 1$, we obtain the synchrotron formula:

$$J_{\text{syn}} = \frac{1}{\sqrt{3}} \left[2K_{\frac{2}{3}}(\xi) - \int_{\xi}^{\infty} K_{\frac{1}{3}}(\lambda) d\lambda \right] \quad (3)$$

where $K_n(x)$ is the modified Bessel functions.

If the trajectory of the electron is sufficiently a circular path at the radiation point, then the formula holds. However, if the trajectory becomes like a straight path, the circular path approximation does not hold because the neglected terms become larger.

【Standard th-trajectory】

A model trajectory for an arbitrary motion where at $t \rightarrow \pm\infty$ the velocity becomes constant while at $t = 0$ the acceleration is maximum (Khokonov and Nitta, Phys.Rev.Lett. **89** 094801 (2002)):

$$\beta_{\perp}(t) = b_0 + b \tanh\left(\frac{t}{T}\right)$$
$$b_0 = \frac{\beta_{\perp 1} + \beta_{\perp 2}}{2}, \quad b = \frac{\beta_{\perp 2} - \beta_{\perp 1}}{2} \quad (4)$$
$$\begin{cases} \beta_{\perp -} \equiv \beta_{\perp}(t \rightarrow -\infty) \\ \beta_{\perp +} \equiv \beta_{\perp}(t \rightarrow +\infty) \end{cases}$$

where T is the interaction time.

With this “th-trajectory” we obtain an analytic expression of radiation spectrum which has two parameters. From (2) and (4), we obtain,

$$J_{\text{th}} = \int_0^{\infty} (1 + 2\nu^2 \tanh^2 z) \sin \Delta_s \frac{dz}{z} - \frac{\pi}{2}$$
$$\Delta_s = \frac{3}{2} \xi \nu [(1 + \nu^2)z - \nu^2 \tanh z] \quad (5)$$

$$\nu \equiv \gamma b \quad (\text{non-dipole parameter})$$

【Quantum Correction】

The well known Baier-Katkov formula has been used for quantum correction. We have

$$\frac{dN}{d\eta} = \frac{\alpha c}{\pi \lambda_c \gamma} J \quad (6)$$

where λ_c is the Compton wavelength, $\eta = \hbar\omega/E$. J is given as

Synchrotron approximation

$$J_{\text{syn}} = \frac{1}{\sqrt{3}} \left[\left(1 - \eta + \frac{1}{1 - \eta} \right) K_{\frac{2}{3}}(\xi^*) - \int_{\xi^*}^{\infty} K_{\frac{1}{3}}(\lambda) d\lambda \right] \quad (7)$$

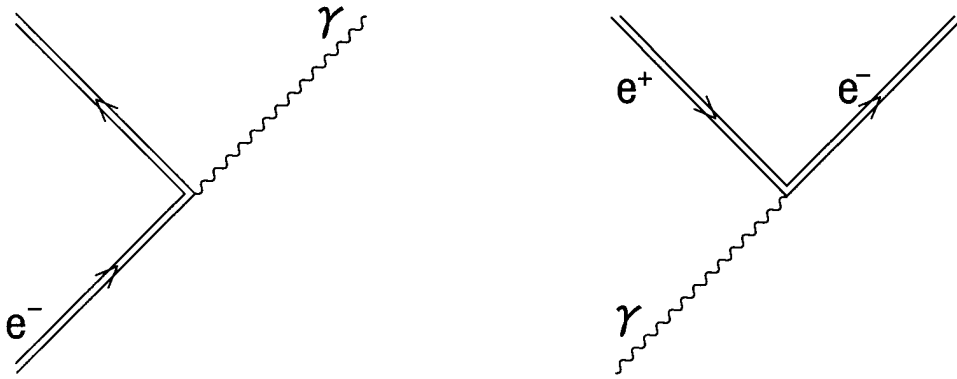
th-trajectory approximation

$$J_{\text{th}} = \int_0^{\infty} \left[1 - \nu^2 \left(1 - \eta + \frac{1}{1 - \eta} \right) \tanh^2 z \right] \sin \Delta_s \frac{dz}{z} - \frac{\pi}{2}$$
$$\Delta_s = \frac{3}{2} \xi^* \nu \left[(1 + \nu^2) z - \nu^2 \tanh z \right] \quad (8)$$

$$\begin{cases} \xi^* = 2\eta/3(1 - \eta)\chi \\ \chi = \gamma\lambda_c F/mc^2 \end{cases} \quad (9)$$

where F is the transverse force from the field. We call χ invariant field parameter.

【Crossing Symmetry】



Once an expression of radiation is established, the pair production can be calculated by using the *crossing symmetry* of the matrix element.

By changing the variables as

$$\begin{cases} E & \text{(initial energy)} & \rightarrow & -E_p & \text{(produced positron)} \\ E' & \text{(final energy)} & \rightarrow & E_e & \text{(produced electron)} \\ \omega & \text{(emitted photon)} & \rightarrow & -\omega & \text{(absorbed photon)} \end{cases}$$

and multiplying the rate of the density of final states to dN

$$\frac{E_p^2 dE_p}{(\hbar\omega)^2 d(\hbar\omega)}$$

we obtain the pair production probability for the trajectory approximation as well as the synchrotron approximation.

【Pair Production】

The pair production probability per unit time is given as

$$\frac{dN}{d\eta} = \frac{\alpha c m c^2}{\pi \lambda_c \hbar \omega} J^{\text{pp}} \quad (10)$$

where $\eta = E^{(+)} / \hbar \omega$ (the energy ratio of the produced positron and the incident photon).

Uniform Field (Synchrotron) Approximation

$$J_{\text{syn}}^{\text{pp}} = \frac{1}{\sqrt{3}} \left[\left(\frac{1-\eta}{\eta} + \frac{\eta}{1-\eta} \right) K_{\frac{2}{3}}(\xi^*) + \int_0^\infty K_{\frac{1}{3}}(\lambda) d\lambda \right] \quad (11)$$

Th-trajectory Approximation

$$J_{\text{th}}^{\text{pp}} = \int_0^\infty \left[\nu^2 \left(\frac{1-\eta}{\eta} + \frac{\eta}{1-\eta} \right) \tanh^2 z - 1 \right] \sin \Delta_s \frac{dz}{z} + \frac{\pi}{2}$$
$$\Delta_s = \frac{3}{2} \xi^* \nu \left[(1 + \nu^2) z - \nu^2 \tanh z \right] \quad (12)$$

$$\xi^* = \frac{2}{3\eta(1-\eta)\chi}$$

non-dipole parameter

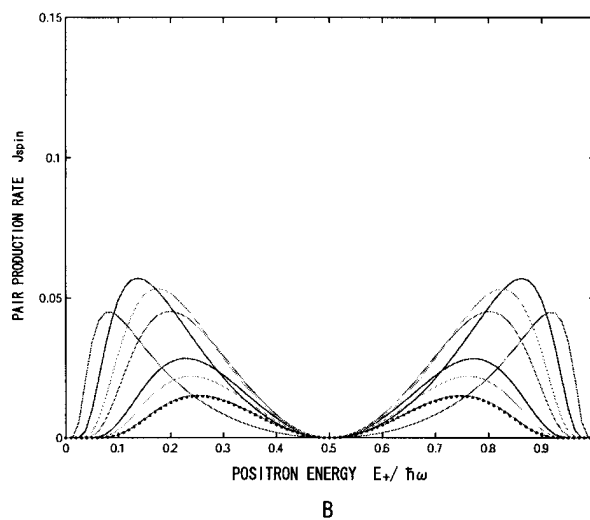
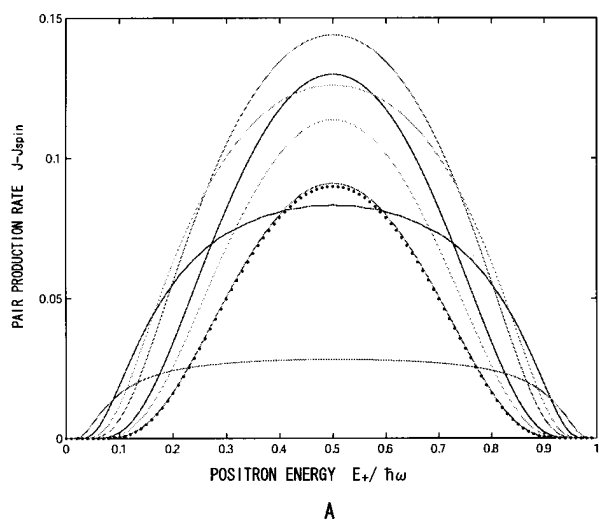
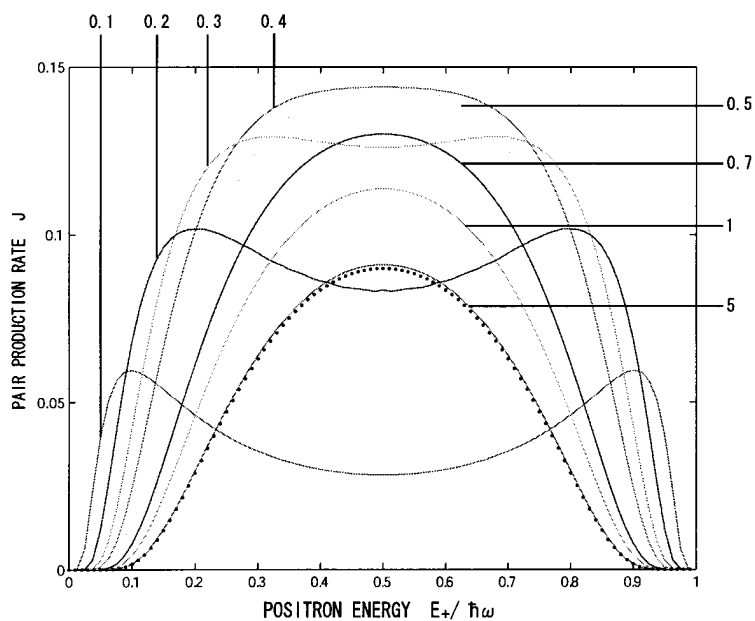
$$\nu = \gamma_p b \quad \gamma_p : (\text{positron's Lorentz factor})$$

invariant field parameter

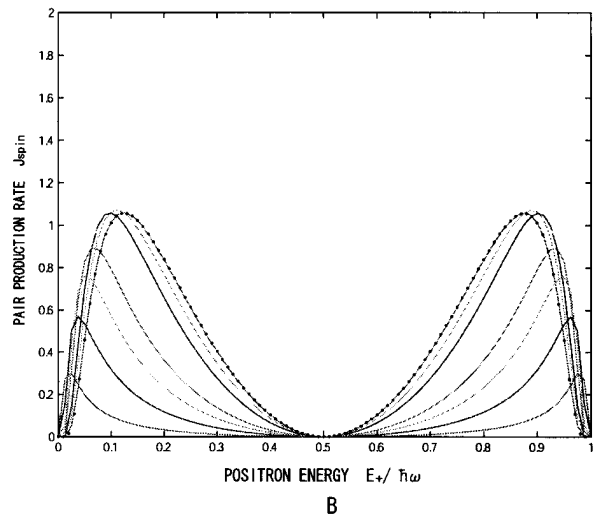
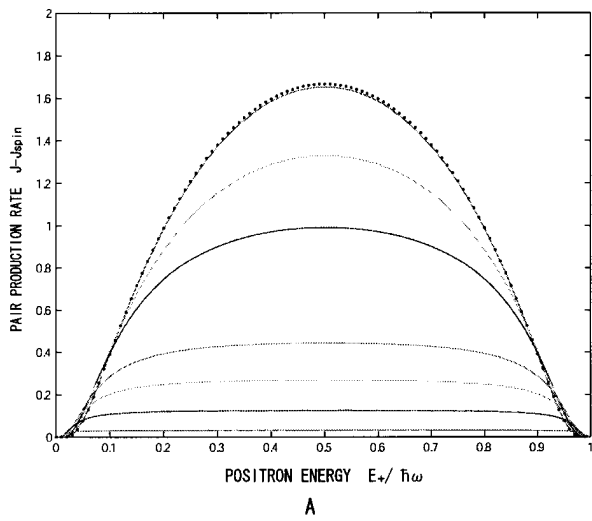
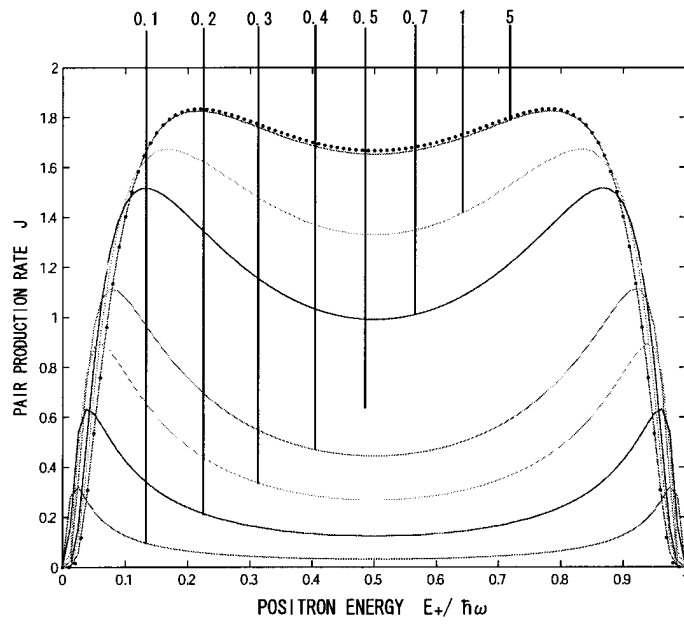
$$\chi = \frac{\hbar \omega \lambda_c F}{m^2 c^4}$$

Pair production probability for $\chi = 1$. dotted line:UFA.

A: J_c . B: J_{spin} (spin-flip contribution)



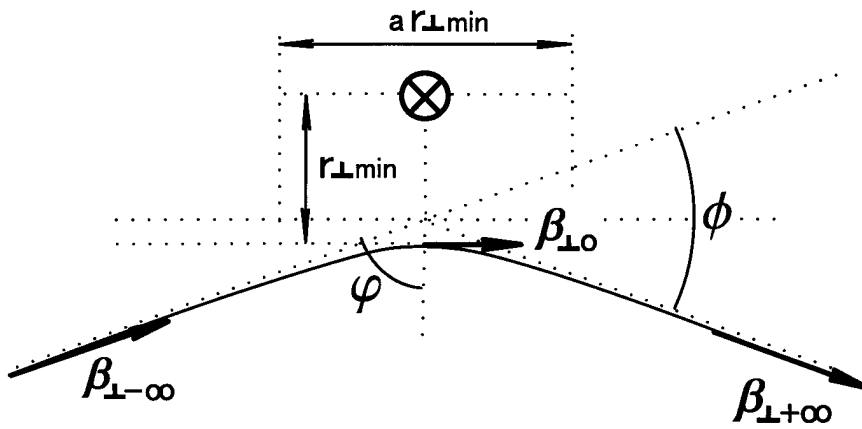
$\chi = 5:$



【Impact Approximation】

non-dipole parameter ν is determined by the scattering angle.

$$\nu = \gamma_p \left| \frac{\vec{\beta}_{\perp+\infty} - \vec{\beta}_{\perp-\infty}}{2} \right| \quad (13)$$



For simplicity, we employ the impact approximation:

$$\gamma_p m c \beta_{\perp+\infty} - \gamma_p m c \beta_{\perp-\infty} = F(\rho_0) \Delta t$$

where we have taken Δt as

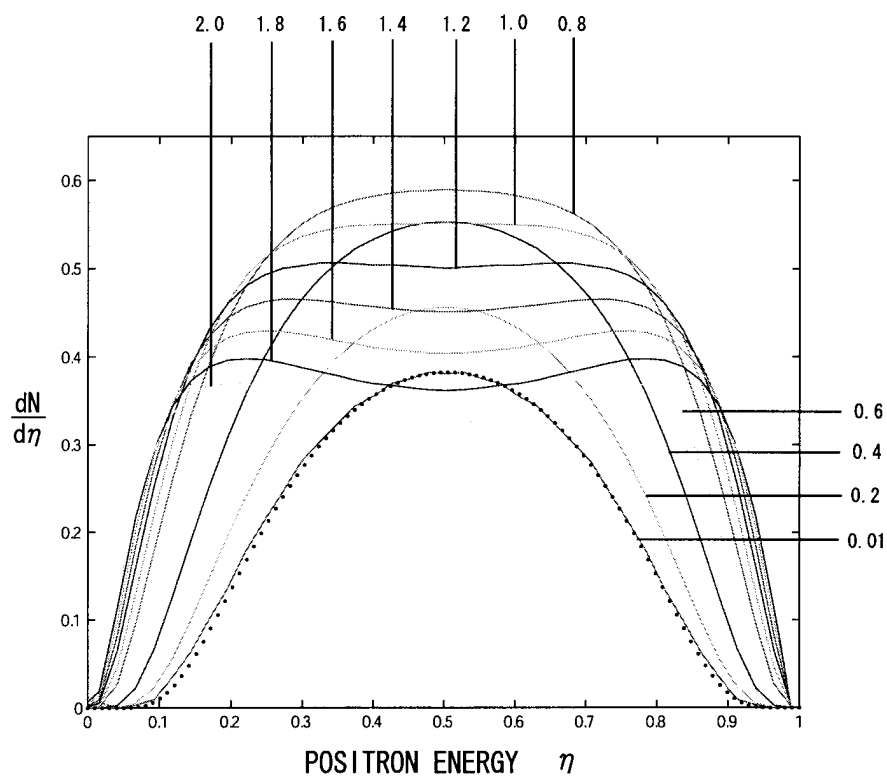
$$\Delta t = \frac{a \rho_0}{c \beta_{\perp 0}}$$

a is the fitting parameter: $a \sim 1$. By taking $\rho_0 = r_{\perp \min}$, the non-dipole parameter becomes

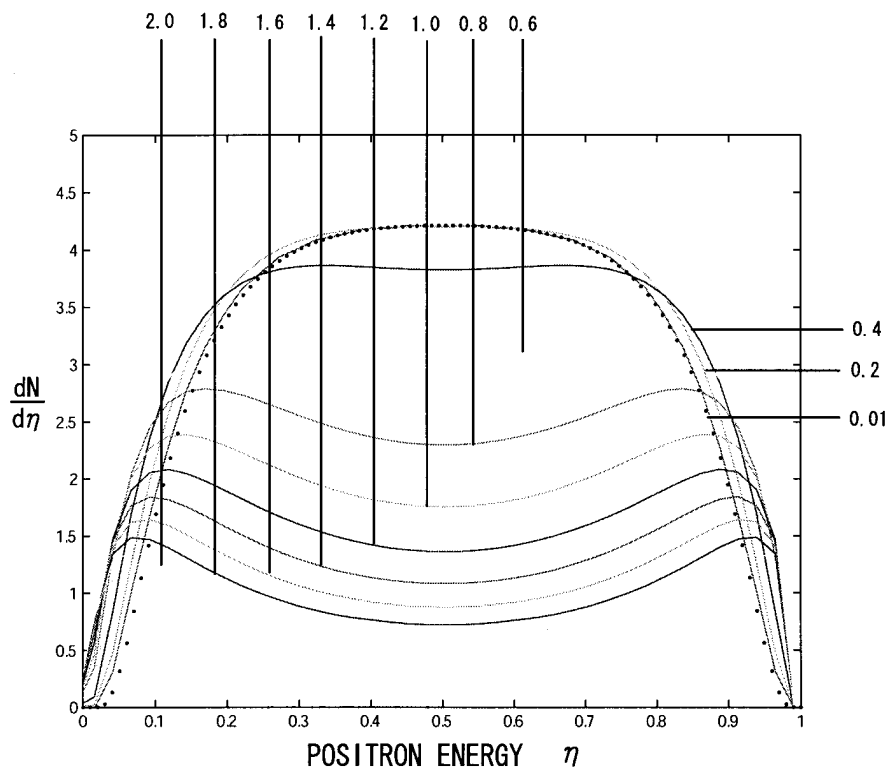
$$\nu = \frac{a F(r_{\perp \min}) r_{\perp \min}}{2 m c^2 \theta_\gamma} \quad (14)$$

where θ_γ is the incident angle of the photon to the crystal axis.

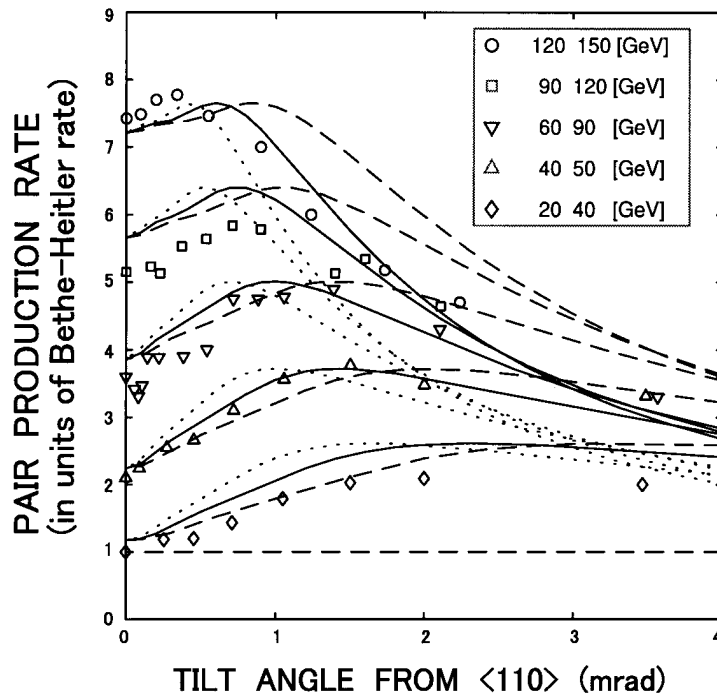
Pair production spectra. The numbers represent the incident angle [mrad]. The dotted line shows UFA. 50[GeV] photon \rightarrow Si< 110 >:



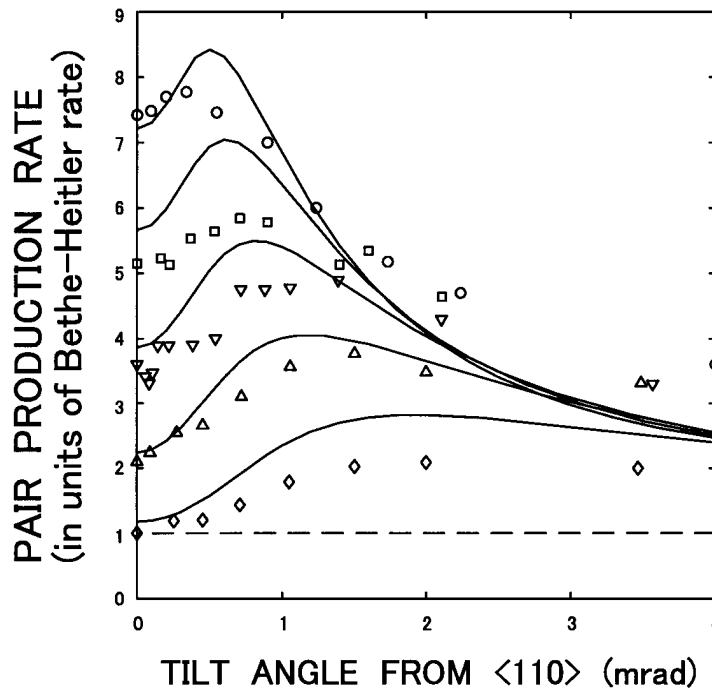
150[GeV]:



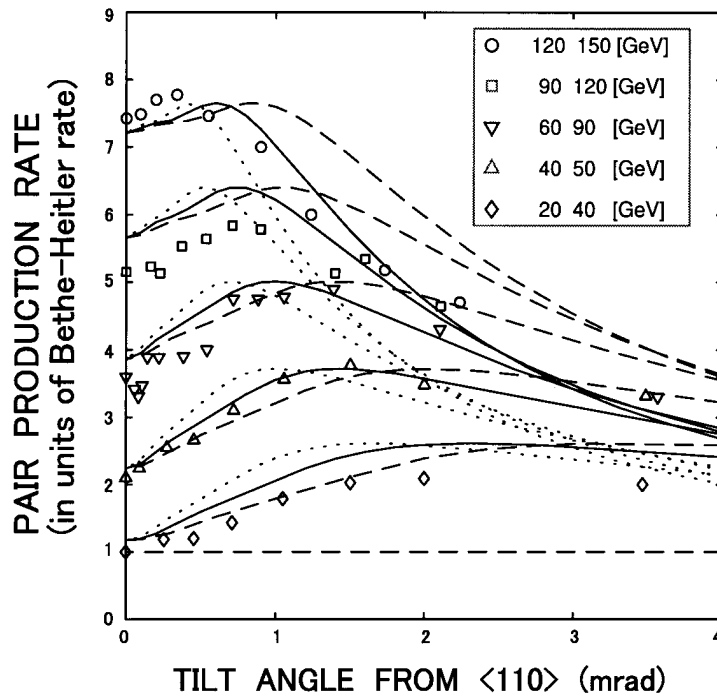
Impulse approximation:



One-string Molière potential ($T = 100\text{K}$):



Impulse approximation:



One-string Molière potential ($T = 100\text{K}$):

